

**NEW INDIVIDUALS FROM THE ALMAHATA SITTA STREWN FIELD: OLD FRIENDS AND BRAND-NEW FELLOWS.** A. Bischoff<sup>1</sup>, S. Ebert<sup>1</sup>, M. Patzek<sup>1</sup>, M. Horstmann<sup>1</sup>, A. Pack<sup>2</sup>, J.-A. Barrat<sup>3</sup>, and S. Decker<sup>4</sup>. <sup>1</sup>Institut für Planetologie, Wilhelm-Klemm-Str. 10, 48149 Münster, Germany. E-mail: bischoa@uni-muenster.de. <sup>2</sup>Uni Göttingen, Geowissenschaftliches Zentrum, Goldschmidtstr. 1, 37077 Göttingen, Germany. <sup>3</sup>Université de Bretagne Occidentale, 29280 Plouzané, France. <sup>4</sup>Meteorite-Museum, Oberstr. 10a, 55430 Oberwesel, Germany.

**Introduction:** After asteroid 2008 TC<sub>3</sub> impacted Earth in 2008, a highly diverse range of meteorite types was identified among the numerous meteorite fragments collected in the Almahata Sitta strewn field, including various types of ureilitic rocks and chondrites (e.g. [1-9]). The petrography and mineralogy of nine new samples were studied here (MS-MU-012 – MS-MU-020), complemented with oxygen isotope compositions obtained for several of the new samples by IR laser fluorination.

**Results:** MS-MU-012 and -019 are the most outstanding new samples. MS-MU-012 is an unbrecciated, ureilitic feldspar-olivine-pyroxene rock. The plagioclase-rich areas do not occur as isolated fragments as observed in (polymict) ureilitic breccias, but appear to represent primary rock constituents. These plagioclase-rich areas display very low REE abundances (<0.15 x CI), except for Eu (~2 ppm; Eu/Eu\* = 250-300). Oxygen isotopes ( $\delta^{17}\text{O} = 2.63\%$ ,  $\delta^{18}\text{O} = 6.98\%$ ; average of 3 measurements) support the relation of the plagioclase-rich aliquot of MS-MU-012 to ureilites. MS-MU-019 is an unusual enstatite- and metal-rich achondrite with two different coexisting enstatite populations (~En<sub>98.5</sub>Wo<sub>1.3</sub> and ~En<sub>96.5</sub>Wo<sub>3.2</sub>). Based on texture, mineralogy, and O-isotopes it might represent a unique type of meteorite. Preliminary results for MS-MU-019 indicate variable  $\delta^{17}\text{O}$  (3.5-4.1‰) and  $\delta^{18}\text{O}$  (6.6-7.5‰) values, which may imply a relationship to E-chondrites or aubrites.

MS-MU-013 has a chondritic texture. Based on the compositions of olivine (~Fa<sub>17</sub>) and pyroxene (~Fs<sub>16</sub>) the rock has to be classified as a H5 chondrite, although the O-isotope compositions ( $\delta^{17}\text{O} = 3.38\%$ ,  $\delta^{18}\text{O} = 4.92\%$ ; mean of 2 analyses) are more related to L/LL. Based on the presence of alabandite and the recrystallized texture, MS-MU-015 clearly is an EL6 chondrite (Fs<sub><0.3</sub>).

MS-MU-014, -016, -017, and -020 are coarse-grained ureilites. The olivine cores in MS-MU-014 have ~Fa<sub>20</sub> and the pyroxene has ~Fs<sub>17</sub>. MS-MU-016 has abundant pyroxene, and the olivine and pyroxene cores have ~Fa<sub>12.5</sub> and ~Fs<sub>11</sub>, respectively. MS-MU-017 is also rich in pyroxene. The olivine and low-Ca pyroxene cores have ~Fa<sub>13</sub> and ~Fs<sub>11</sub>, respectively. Frequently observed Ca-pyroxenes have variable compositions (~Fs<sub>6-36</sub>Wo<sub>30-39</sub>). The olivine in MS-MU-020 is strongly zoned with the highest Fa content of ~Fa<sub>21</sub>; low-Ca pyroxene cores have ~Fs<sub>18.5</sub>. MS-MU-018 is a heavily-shocked, fine-grained ureilite with abundant opaque phases (metal, sulfides).

**References** [1] Jenniskens P. et al. 2009. *Nature* 458:485-488. [2] Bischoff A. et al. 2010. *Meteoritics & Planet. Sci.* 45:1638-1656. [3] Horstmann M. and Bischoff A. 2014. *Chemie der Erde - Geochemistry* 74:149-184. [4] Goodrich C. A. et al. 2014. *Elements* 10:31-37 [5] Bischoff A. et al. 2012. *Meteoritics & Planet. Sci.* 47:A71. [6] Horstmann M. et al. 2010. *Meteoritics & Planet. Sci.* 45:1657-1667. [7] Zolensky M. E. et al. 2010. *Meteoritics & Planet. Sci.* 45:1618-1637. [8] Horstmann M. et al. 2012. *Meteoritics & Planet. Sci.* 47:A193. [9] Bischoff A. et al. 2014. *Proc. Natl. Acad. Sci.* 111:12689-12692.